



Vera C. Rubin Observatory  
Data Management

# Calibration Generation, Verification, Acceptance, and Certification

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DMTN-222

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## Abstract

This technote defines the best practices to be used for calibration generation, the verification that that calibration meets requirements, and when deciding if the calibration should be accepted for use in processing at both the summit and USDF. Other aspects of calibration product management, including the contents of the appropriate “defaults” collection, are also discussed.

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3	2024-04-17	Updated draft reflecting processes actually in use.	Chris Waters
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7	2025-06-09	Add TAXICAB tracking and photodiode ingest instructions.	Chris Waters

*Document source location:* <https://github.com/lstt-dm/dmtn-222>

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# Calibration Generation, Verification, Acceptance, and Certification

## 1 Introduction

The purpose of this technote is to provide guidance on the procedures that are used for the construction and management of calibrations. These guidelines shall be followed for any calibration that will be added to the main butler repositories. For the purposes of this document, we will consider three cases of calibrations.

- Calibrations generated for widespread use, using the main butler repository. These are referred to as “combined calibrations” below, and indicate the calibrations that are used for science processing.
- Curated calibrations that are defined by an `obs_` package and must be ingested to the butler repository, as they cannot be directly generated from raw data. The camera geometry calibration is an example of this type of calibration.
- Calibrations that have been exported from one butler repository for use in another.

Additional private calibrations produced for tests may also exist, but as those will only exist in a user-space collections, they will not be discussed further in this document.

Briefly, calibration construction involves the following steps:

**Generation** An appropriate set of exposures is chosen and processed through the correct `cp_pipe` pipeline.

**Verification** The proposed calibration is used to process exposures through the matching `cp_verify` pipeline. This processing measures a set of quality metrics, as defined by DMTN-101, to determine if the newly made calibration meets requirements.

**Certification** The proposed calibration is certified for a particular usage date range. These are generally open ended, with only the start date defined. We expect to know the start date for the majority of all calibrations, as they should correspond with changes to the camera.

**Approval** The TAXICAB considers the proposed calibrations and their associated verification results, and makes the decision on whether the proposal is accepted for use.

**Distribution** The collection containing the new calibrations are included in the main calibration collection chain, for all repositories that need the updated calibration.

Figure 1 displays the relationship between the various stages of construction, validation, and use of combined calibrations. **CZW: Is this still useful?**

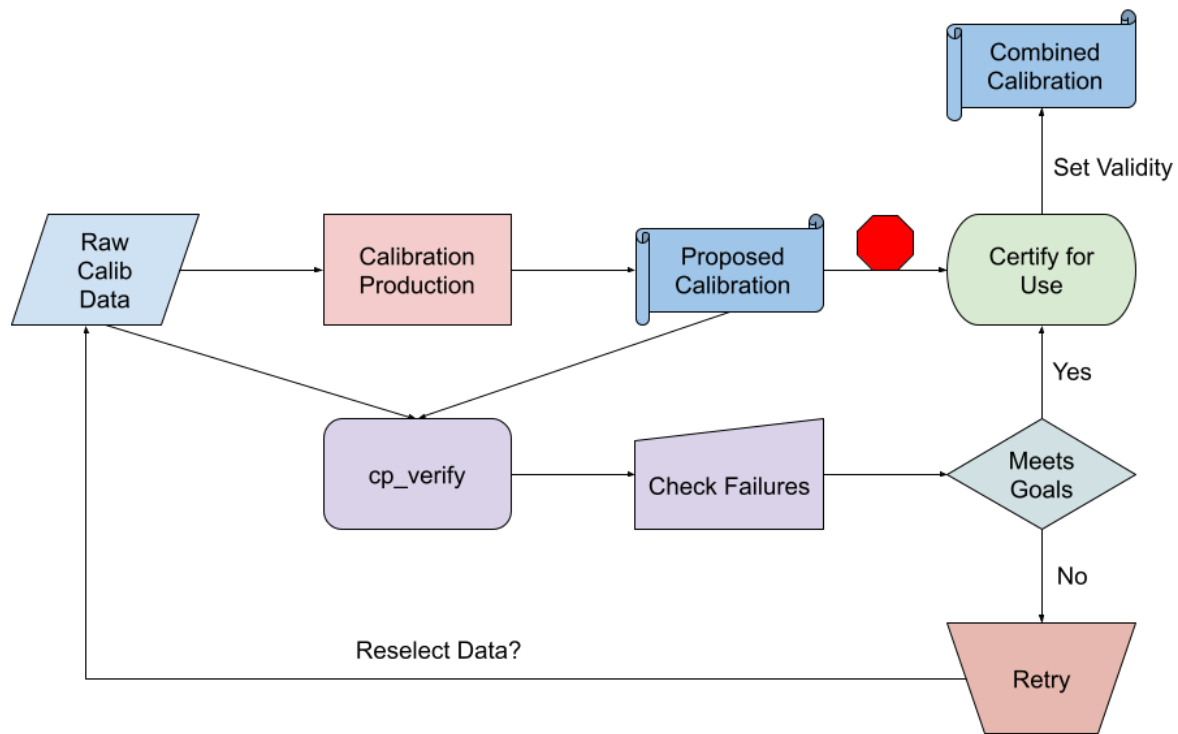


FIGURE 1: Flowchart of the calibration construction process.

## 2 Collection naming

Consistent collection names make the management of calibrations easier. JIRA tickets are used to ensure that these collection names are unique, and that there is a clear location to find

the construction artifacts for later analysis. In addition to this ticket, a short string explaining the purpose of the calibration set should be included in the collection name to provide a human readable “tag.” The following collection name patterns, based on the recommendations in DMTN-167 should be followed for all calibrations that will be approved by the TAXICAB.

The calibration generation should use the form

```
$INSTRUMENT/calib/$TICKET/$TAG/${CALIB_TYPE}Gen.${RERUN_ITERATION}
```

where \$INSTRUMENT is the camera name, \$TICKET is the JIRA ticket value, \$TAG is the short human readable string, \$CALIB\_TYPE is the calibration type being generated, and \$RERUN\_ITERATION is a date string of the form YYYYMMDDv indicating when the calibration was made, with a trailing character to be incremented if the generation must be retried. As an example, a hypothetical new bias would have a collection name like LATISS/calib/DM-12345/voltageChange/biasGen.20220915a.

For verification, a similar form is used:

```
$INSTRUMENT/calib/$TICKET/$TAG/verify${CALIB_TYPE}.${RERUN_ITERATION}
```

with the same elements as for generation. The certification process (see below) also creates a new CALIBRATION collection, which should just contain the calibration type:

```
$INSTRUMENT/calib/$TICKET/$TAG/${CALIB_TYPE}.${CERTIFICATION_RERUN}
```

These CALIBRATION collections should be added to a CHAINED collection that uses the collection base only up to the ticket. This ticket-level CHAINED collection provides a way to add and remove all of the calibrations constructed as part of that ticket in one operation. In addition, by ensuring that new calibration collections are prepended to the top level CHAINED collection, we can avoid needing to set end dates during calibration certification. The butler search for calibrations completes when the first matching calibration is found. So, as an example, by putting a collection with a start date of 2024-10-01 before one with a start date of 2024-06-01, we can ensure that an exposure taken on 2024-10-15 will be processed with that newer calibration, but one taken on 2024-09-01 will be processed with the older one.



## 3 New Combined Calibrations Construction

A record of the calibration construction process should be retained and attached to the JIRA ticket managing the work, with all commands executed and exposure selections recorded. Having this record will allow for understanding what happened during construction, in case the final products have problems.

### 3.1 Generation

Combined calibrations will be generated directly from raw exposures as much as possible. The tasks and pipelines in the `cp_pipe` package can produce all of the calibrations that are currently used for image processing, and can be supplemented as new corrections are developed. The main documentation for calibration construction is included in `cp_pipe` at <https://pipelines.lsst.io/v/daily/modules/lsst.cp.pipe/constructing-calibrations.html>, but the main points will be summarized here.

Calibrations are inter-dependent, and so the construction of one type may require precursor calibrations to be built first. Figure 2 shows the current dependence, with each box pointing to the calibrations that they depend on. The result of this is that changes in one calibration (such as the gains derived from the photon transfer curve) require other calibrations (the linearity, the brighter-fatter kernel, and the charge transfer inefficiency) to be built as well.

The `observation_type` and `observation_reason` of the input exposures should match the calibration type to be constructed, with the exception of the fringe and crosstalk calibrations, which are constructed from science exposures. Most calibrations can be constructed from a single set of daily calibrations, with the number of bias, dark, and flat frames in these sets (generally of order 15-20) sufficient to create a usable combined calibration. Dense PTC curves will require many more inputs (on the order of 100 pairs of exposures), and we currently expect that we will have dedicated observation sequences for this purpose.

Calibrations constructed for general use should be able to use the version of the `cp_pipe` tasks and pipelines on the main github branch. It is preferable to keep code development separate from the calibration construction, but it is expected that these will likely be coupled during commissioning.

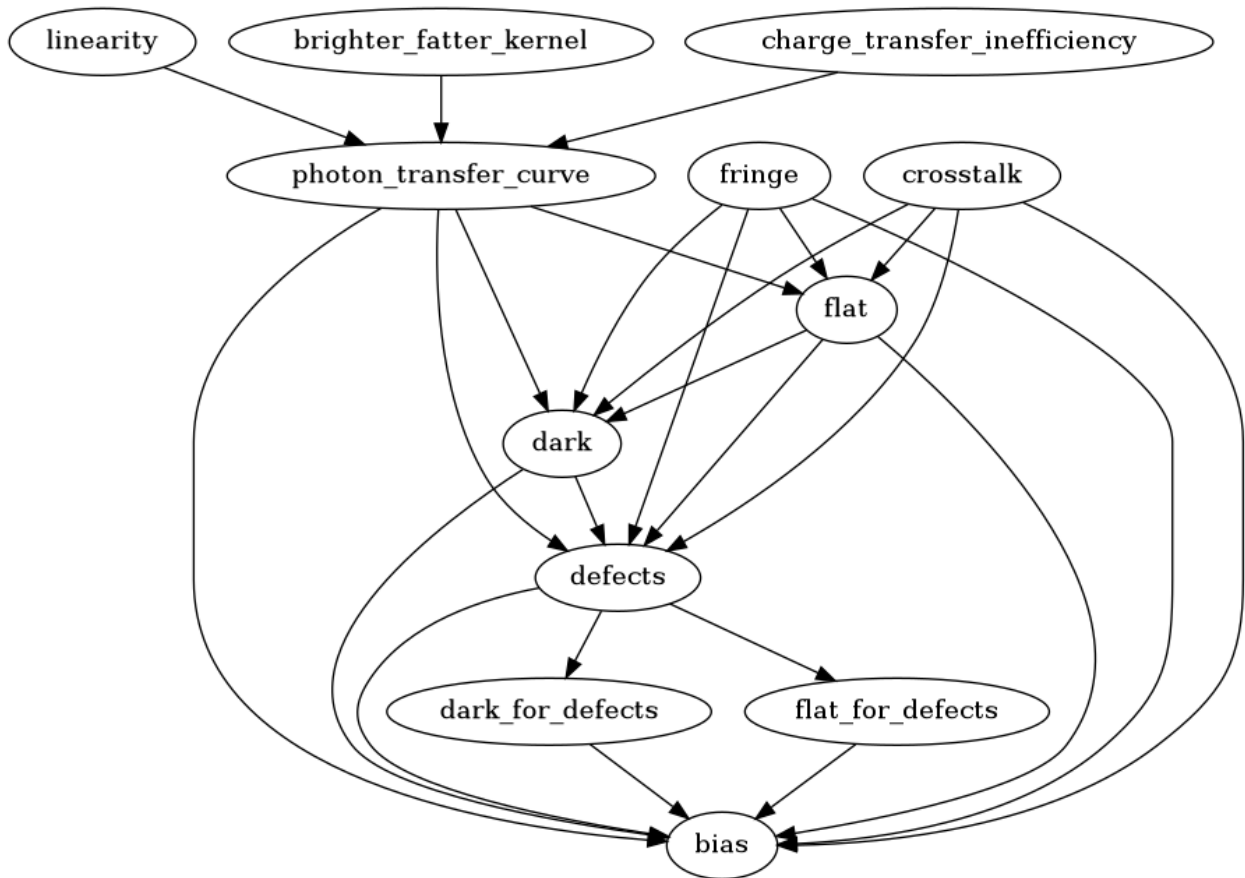


FIGURE 2: Dependency charge of calibration products. The arrow indicates the parent calibration.

To ensure all butler repositories have a consistent set of calibrations, we have decided that only one processing location should perform the calibration construction steps. The US Data Facility (USDF) is now operational, all calibrations used for the survey will be generated there. The process for transferring the calibrations to other butler sites is discussed below in Section 3.6.

## 3.2 LSSTCam Flats

At the present time, we have two ways of generating “white light” dome flats with LSSTCam data. In the future, we will be adding monochromatic flats from scans with the tunable laser.

The term “white light” refers to illumination with the calibration system LED projector, with broad (“white”) and not narrow (laser) illumination. The light itself is not white, and there are 10 LEDs covering the wavelength range of the ugrizy filter system.

The original configuration (also used for the daily flats) can illuminate each of the griz filters with 2 LEDs at a time. These can be processed with the `cpFlat.yaml` pipeline and the `bps_flat_[band].yaml` templates.

The new configuration produces two sets of flats for each filter, one with a bluer LED and one with a redder LED (each appropriate for the filter). These red and blue flats then get individually processed and weighted and combined at the end. We call these “anaglyph” flats, in an analogy with the anaglyph stereoscopic photographic technique where two red/blue images are overlaid and viewed with special red/blue glasses. These can be processed with the `cpFlatAnaglyph.yaml` pipeline and the `bps_flat_anaglyph_[band].yaml` templates.

## 3.3 Verification

A verification report is then generated from the generation and verification collections, and will be supplied to the Telescope And auxiliary Instrumentation Calibration Acceptance Board (TAXICAB), as defined in RTN-075.

Once the proposed calibrations have been generated, the calibration should be used for processing using the `cp_verify` tasks and pipelines. These tasks measure quality metrics from those processed exposures, and identify any test failures. At a minimum, the exposures

used to construct the calibration should be included, as this can identify problematic inputs that degrade the calibration quality. An example of this is saturated flat exposures, which do not flat-field well, and should not be included in the final flat calibration. In running the `cp_verify` tasks, the input butler collections specified should have the construction RUN collection placed at the beginning of the list, to ensure that the verification process will find and use the calibration we wish to verify.

Exposures from outside the set used for construction should be added to provide insight into the expected validity range for the calibration. For calibrations generated to correct for known camera changes (such as a sequencer upgrade), we will likely need to generate the combined calibration from a limited set of inputs, and will therefore rely on `cp_verify` metrics to monitor those calibrations. In cases where we have multiple days of observations that can be used, including exposures from multiple dates in the verification set can be used to As long as the metrics on those exposures remain within the limits defined in DMTN-101, the calibration may continue to be valid for the date range including those additional exposures. This can be used to establish the valid date ranges to be used when certifying the calibration.

The `cp_verify` pipelines will generate and publish `analysis_tools` “core” metrics and plots to cover the DMTN-101 tests. These metrics and plots will also include useful diagnostic results based on the camera team `eo_pipe` tests. Further “extended” metrics and plots may also need to be generated to supply additional debugging information about the calibrations.

The `cp_verify` report is generated by a command such as

```
$CP_VERIFY_DIR/bin/cpv_report.py -r embargo \  
    -O ~/public_html/cpv_reports/TAXICAB-15 \  
    -c LSSTComCam/calib/DM-47447/gainFixup/verifyFlat-g.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/verifyFlat-r.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/verifyFlat-i.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/verifyBias.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/verifyDark.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/flatGen-g.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/flatGen-i.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/flatGen-r.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/darkGen.20241107a \  
    -c LSSTComCam/calib/DM-47447/gainFixup/atoolsDark.20241108a \  
    \
```

-c LSSTComCam/calib/DM-47447/gainFixup/atoolsDarkDet.20241108a

As many collections and types of calibrations can be added to the list of inputs. All applicable verify collections should be included, and although the generation collections should be linked from the verify collection (allowing them to be searched for data products), it is generally safe to add them to the list as well. This report constructs set of HTML documents and web-friendly images so the calibration quality can be checked. The report is known to be incomplete, as there is a large backlog of tickets for adding plots and metrics to the code.

### 3.4 Certification

Once the new combined calibration has been generated and verified, it can be certified for use for a given date range. Calibrations that have been constructed due to a camera or telescope change, or that are being built to replace another calibration that is no longer within the test specifications, should always have a starting validity date, with the end date left open. This ensures that future data taken will always have valid calibrations for processing.

If historical calibrations are being constructed, the end date should be known from the daily calibration processing results stored in the visit database (see below). Future development is needed to allow calibrations to be recertified to update the date ranges.

The certification process can be done with a simple butler command:

```
butler certify-calibrations $REPO \  
  --begin-date $START_DATE \  
  $INSTRUMENT/calib/$TICKET/$TAG/biasGen.$RERUN \  
  $INSTRUMENT/calib/$TICKET/$TAG/bias.${CERT_RERUN} \  
  bias
```

where the `$START_DATE` is the ISO-8601 datetime in TAI coordinates, and `$CERT_RERUN` is a dated rerun string as used above (as the generation and certification may take place on different dates).

Note that certification is purely a “database operation,” with the certified collection gaining a “link” to the dataset in the construction collection. This means that if the underlying generation

collection is removed, the calibrations may be lost as well.

### 3.5 Approval

With the calibrations built, verified, and certified, a TAXICAB “hailing” ticket should be created, with a verification report attached for consideration. Any member of the Rubin Observatory Team is welcome to join the TAXICAB meeting, and the meeting will make decisions as a consensus. Any additional processing that is suggested by the TAXICAB should be defined and run prior to the TAXICAB meeting, which will have a planned weekly time slot. Currently, this meeting is scheduled for Tuesday at 2:30pm Project Time. If no open TAXICAB hailing tickets exist, this meeting will be skipped. Scheduling a TAXICAB meeting generally indicates that the associated ticket has been marked as “Flagged,” as it is awaiting the decision on acceptance.

The TAXICAB will consider the verification reports, identify any potential issues with the calibration set, and determine if any verification test failures warrant restarting the construction process to address the issues. Ideally, all verification metrics will succeed, and a quick check of residual exposures will show no unexpected features. In the more likely case that some fraction of these tests fail, the TAXICAB will be tasked with deciding if the failures are fatal and that the calibration should be fully rejected, or if the failures are small enough in number or impact that the calibration can be accepted for use despite them. The TAXICAB will operate on a consensus basis, to ensure that all stakeholders have input on this process.

If the calibrations were built using a ticket/development branch of any software, those code changes must be reviewed and approved through the standard DM process prior to hailing the TAXICAB. If no new code was added, then the approval of the TAXICAB can be used as the review process to close the initial generation ticket.

Managing the approval for calibrations will follow a three-ticket process. As described above, a construction ticket is used to manage all construction, verification, and certification steps, and is included in all output collection names. A TAXICAB ticket is used to manage the approval of the calibrations, and to contain pointers to the verification report, the generation ticket, and the final ticket, which manages the deployment of these calibrations. This deployment ticket is used for copying the calibration into all necessary public repositories, as described below. If code was added as part of the construction, that work should be reviewed on the construction ticket prior to the TAXICAB meeting. Otherwise, the TAXICAB approval (by marking that ticket “Adopted”) should also mark the construction ticket as reviewed. The deployment ticket will

be self-reviewing in the future: a trivial processing task should be run at each repository that has had the new calibrations deployed, with checks that those new calibrations are correctly selected and used.

A summary of the TAXICAB tickets is currently compiled on the project Confluence <https://rubinobs.atlassian.net/wiki/spaces/DM/pages/281673762>. This page should be updated as new TAXICAB tickets are generated, to ensure that we have a central location to track the various tickets. As some repositories may have the distribution step skipped for various reasons, such as not having data for the camera the calibrations were created for, this page also lists which repositories the calibrations on that ticket were copied to. A central archive of the exported calibrations needs to be established, so that the calibration CHAINED collection can be reconstructed at all data facilities.

## 3.6 Distribution

Upon approval of the TAXICAB, the calibrations can be distributed for use. A separate distribution ticket should be created to handle this work, and linked to both the construction ticket and the TAXICAB ticket.

There are three different distributions that need to be handled. First, calibrations in `main` need to be chained into the default calibration collection. Second, calibrations generated in `main` need to be copied over to the `embargo` repo at USDF if they need to be applied to embargoed images. Third, calibrations need to be copied to the Summit, Base Test Stand, and other Data Facilities via `rucio`. In the latter two cases, after calibrations have been copied they need to be chained into the default calibration collection.

In most cases, when new forward-looking calibrations are deployed as the default, the new CHAINED collection which was created as part of the calibration generation can be prepended to the top level calibration CHAINED collection, installing the calibrations for use by default.

### 3.6.1 Exporting and Importing Calibrations

The calibrations must then be exported for use in other repositories and to create a local archive of calibrations at USDF. An example command to do such is

```
butler export-calibs $REPO \  
  ./export_directory \  
  LATISS/calib/DM-XYZ \  
  LATISS/calib/DM-XYZ/voltageChange/bias [...]
```

This command exports the files into the `export_directory` location, and constructs a YAML description of the calibrations and their collections. Exporting the ticket-level CHAINED collection will export all of the children CALIBRATION collections, making it the preferred way to export calibrations.

This `export_directory` must then be transferred to the location of the new repository, where it can be imported with the command

```
butler import $NEW_REPO --transfer copy \  
  --export-file ./export_directory/export.yaml ./export_directory \  
  -s instrument -s detector -s physical_filter
```

The `--transfer copy` is strongly suggested, as this will copy the files into the repository data-store, removing any dependency on the `export_directory`. The three `-s` arguments indicate that the `instrument`, `detector`, and `physical_filter` definitions contained in the YAML description should be skipped, as they will already exist in a repository that has been set up for the appropriate camera.

The newly imported collections will not by default be part of the public calibration collection. To do so, the new collections must be added to the collection chain. Using the following command with the 'prepend' mode will add the new collections to the start of the collection chain, making them available.

```
butler collection-chain $NEW_REPO --mode=prepend LATISS/calib \  
  LATISS/calib/DM-XYZ \  
  LATISS/calib/DM-ABC
```

The distribution ticket should be able to be self-reviewed, after confirming that at least one exposure from the validity range of the new calibrations can be processed through `IsrTask`,



and that the output processed exposure has the correct calibration information recorded in its header.

The following table lists the current set of facilities, repositories, and which cameras are deployed in that repository.

Data Facility	Repository	Camera
USDF	embargo_old	LATISS
		LSSTComCam
		LSSTComCamSim
		LSSTCam
USDF	embargo	LATISS
		LSSTComCam
		LSSTComCamSim
		LSSTCam
USDF	main	LATISS
		LSSTComCam
		LSSTComCamSim
		LSSTCam
USDF	/repo/ir2	LSSTCam
Summit	LATISS	LATISS
Summit	LSSTComCam	LSSTComCam
		LSSTComCamSim
Summit	LSSTCam	LSSTCam
Tucson Test Stand (TTS)	LATISS	LATISS
Tucson Test Stand (TTS)	LSSTComCam	LSSTComCam
		LSSTComCamSim

### 3.6.2 Distribution With Rucio

The way to copy calibration files from `main` to the Summit, BTS, and the other DFs is via `rucio`. Note that `rucio` does not have access to the `embargo` repo.

Using `rucio` involves first registering the datasets with the `rucio` service, and then creating a “rule” which says where the files should be synchronized. Once the rule is created then the `rucio` service will do the necessary synchronizations, and will place the calibrations in special

S3 calibration buckets at BTS, the Summit, and at any other DFs that are configured. The S3 calibration buckets are configured such that only rucio has write access, and these files cannot be directly deleted (on purpose or accident). We have not yet figured out the plans for removing old calibrations.

At the current time rucio will only synchronize the files, and it is the responsibility of the person doing the deployment to copy over the above `export.yaml` and do the ingestion and the final collection chaining. In the near future we expect the ingestion will be done automatically, though explicit deployment chaining is going to be a manual task for the foreseeable future.

Currently, there is work being done to create automatic Rucio “subscriptions” such that the calibration group merely define a dataset and the automated system handles the distribution. This depends on pre-determined naming conventions. For calibrations to be distributed to the Summit and BTS, the DATASET name should be of the form `Dataset/$INSTRUMENT/calib/$TICKET`, matching the chained calibration collection name. For calibrations that are not to be distributed to the Summit and BTS, and only the other DFs, the DATASET name should be of the form `Dataset/historical/$INSTRUMENT/calib/$TICKET`, matching the chained calibration collection name. These are calibrations that are not to be used for online processing, but only for (e.g.) data releases where the calibrations only apply to older data.

The following is how a set of darks and flats can be synchronized.

```
# This is the collection where calibs will be copied from
COLLECTION=$INSTRUMENT/calib/$TICKET
# This dataset name must be unique
DATASET=Dataset/$INSTRUMENT/calib/$TICKET

# Loop over all the types of calibration products in the
# collection. In the future rucio-register may be able
# take an array of dataset types.
# Be careful with this step, as you need to know all the
# types of calibs in the collection.
for dtype in dark flat; do
    rucio-register data-products \
        -s 10 \
        -C /sdf/data/rubin/shared/calibration_archive/rucio/main-calib-config.yaml \
```

```
-r main \  
-t $dstype \  
-c $COLLECTION \  
-d $DATASET  
done  
  
# Close the dataset so that no further files can be  
# associated with it.  
rucio did update --close ancillary:$DATASET  
  
# List the contents of the dataset to confirm that everything  
# looks as expected.  
rucio did content list ancillary:$DATASET  
  
# Define a rule to synchronize the files at 3 sites (note that  
# the number of copies must match the number of rucio sites).  
RULEID=$(rucio rule add \  
  --copies 3 \  
  --rses 'SLAC_BUTLER_DISK|BASE_CALIB_DISK|SUMMIT_CALIB_DISK' \  
  ancillary:$DATASET \  
  )  
# Show detailed information on the synchronization of the files.  
rucio rule show --rule $RULEID  
  
# Show a concise table of the synchronization state of all the  
# files and locations. This command should be run every once in  
# a while until all the files have completed synchronization.  
rucio replica list dataset ancillary:$DATASET
```

Once the replica list shows that everything is synchronized among the sites, the `export.yaml` can be copied to the BTS and Summit and imported as such. It is very important to get the file prefix correct, as well as to use the `direct` import method which tells the database that the files are already safely in the calibrations bucket.

```
butler import $REPO s3://butler@rubinobs-calibrations/rucio/ancillary --export-file `pwd`/export.yaml
```

After this the calibration collection chain can be updated as with main and embargo.

### 3.6.3 TAXICAB Tracking

For TAXICAB tracking, we have a page at <https://rubinobs.atlassian.net/wiki/spaces/DM/pages/281673762/TAXICAB+Tracking+Page>. The goal of this page is to provide a single list of all TAXICAB tickets, the associated calibration generation and deployment tickets, and the deployment status for various repositories. It is strongly encouraged that this is kept up-to-date as calibrations are deployed, until we have a way to fingerprint deployed calibrations to ensure uniformity across sites.

## 3.7 Auxiliary Data Products

In addition to calibrations that are directly used in the processing of data, there are other data products that are managed and distributed by the calibrations team to ensure that repositories are consistent. These datasets are generally added to the `$INSTRUMENT/default`s CHAINED collection, which links the raw exposure collection, the top level CHAINED calibration collection, and other standard data products that are needed for completely processing. The two most common of these data products are reference catalogs (refcats) and skymaps.

### 3.7.1 Refcats

**CZW:** This needs to be written.

### 3.7.2 Skymaps

The skymaps contain the spatial information about the tracts and patches used for coadd construction. The standard skymap configuration files can be found in **CZW: a TBD archive**. The information about these skymaps is stored in the butler database, and not as standard on-disk files, so care is needed to ensure that the database has enough storage space for the 3-5GB used by each skymap.

Registering the skymap is easy when using a predefined configuration:

```
butler register-skymap $REPO \  
    -C skymap-lsst_cells_v1.config
```

As these are stored in the database, a butler repository can have the available skymaps checked by running

```
butler query-dimension-records $REPO skymap
```

### 3.7.3 Other Data Products

Currently, FGCM lookup tables and “pretrained models” are the other data products that are connected to the default collection. **CZW: This needs written as well.**

## 3.8 Calibration Construction Checklist

- File DM ticket for calibration construction that will hold the details of that process
- Select inputs for calibration
- Run generation pipeline from `cp_pipe`.
- Run associated verification pipeline from `cp_verify` on the same inputs, supplemented with similar exposures to test for time variability.
- Create verification report.
- File TAXICAB ticket with verification report and any other information.
- File deployment ticket.
- Upon approval, export/import the calibrations to all other repos. This may be further automated in the future.
- Ensure summit calibrations are tagged correctly.

## 4 Daily Calibrations

Daily calibrations are used to monitor the camera and telescope for changes. The daily calibration processing will simply verify these newly taken exposures against the existing calibration set as shown in Figure 3. This allows the long-term stability of the calibrations to be monitored. We do not plan to ever generate new combined calibrations automatically from the daily calibration scripts.

The `cp_verify` pipeline from the daily processing should run using the “`butler+sasquatch`” butler repos, which ensure that the `analysis_tools` metrics are dispatched to the summit Chronograf (<https://summit-lsp.lsst.codes/chronograf>). There will be a new dashboard available to display these metrics as they change from day-to-day, providing a way to monitor camera/system changes.

The verification results from the daily calibration processing will eventually issue LOVE-based alarms if any tests fail. This should notify the calibration team members and result in an investigation to determine if updated calibrations need to be produced.

## 5 Curated Calibrations

Curated calibrations are those calibrations that cannot easily be generated from a series of exposures, or that require special hardware that will not be available at the summit. Currently, the camera geometry calibration is the most important curated calibration in wide use. These calibrations are ingested via the `butler write-curated-calibrations` command. This command by default will attempt to write to the main `$INSTRUMENT/calib` collection. This is generally not desired, as it is useful for that collection name to point to a CHAINED butler collection, to allow for the calibration management process described above. Instead, a JIRA ticketed collection name should be used, as the following example illustrates for the LATISS camera:

```
butler write-curated-calibrations $REPO lsst.obs.lsst.Latiss \  
  --collection LATISS/calib/DM-XYZ --label DM-XYZ
```

This will ensure that the calibrations can be chained into the main collection as detailed above.

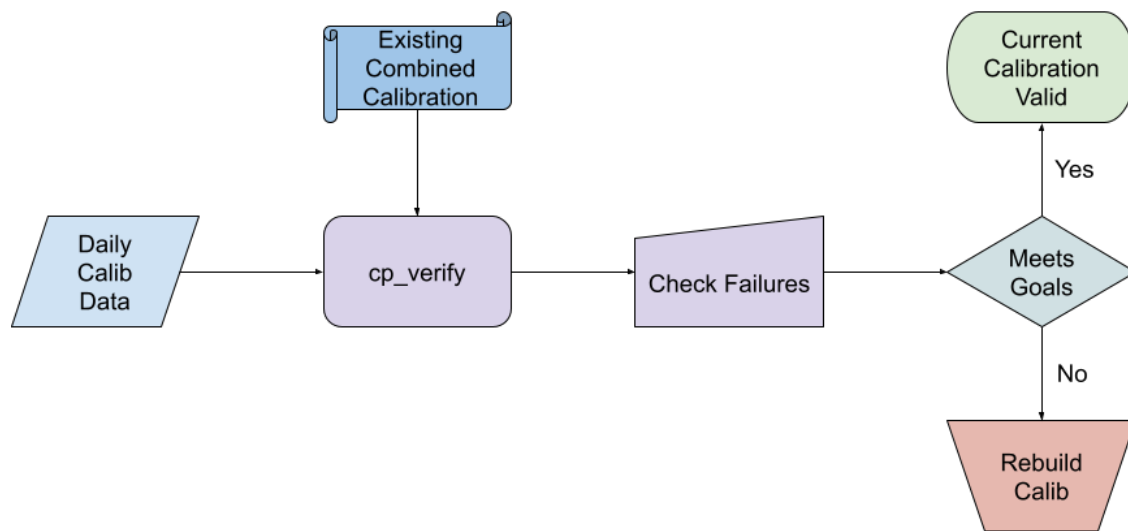


FIGURE 3: Flowchart of the daily calibration process.

## 5.1 Semi-curated Calibrations

We have a number of auxiliary instruments that generate “calibration-like” data products that are useful to provide independent measurements to support the activities of the main cameras. For these to be easily usable, they must be ingested into the butler with a unique dataId corresponding to a single exposure. This matching is performed using the instrument name and group field, and each exposure can only have one associated calibration per dataset type. These files are modified during ingest to add additional header information needed to support direct reads from the butler, and so until that information is added to the original files on write, they will need to be ingested with the `--transfer=copy` argument. This transfer mode is enforced by the task’s configuration validation step, to prevent unreadable files from being ingested.

### 5.1.1 Photodiode Measurements

The photodiode measurements are the first of these calibration-like data products. These data are also referred to as “electrometer” data, as the output comes from the electrometer that measures the current coming out of a photodiode that is in the same optical path as the camera, providing an independent flux calibration for their matching exposures. The original files are written to the Large File Annex (LFA), so the instructions for dealing with S3 data (available at <https://rubinobs.atlassian.net/wiki/spaces/LSST0ps/pages/45636680/USDF+S3+Bucket+Organization>) are useful for checking what data exists. The two main ways to do these checks are a simple `ls`:

```
mc ls lfa/rubinobs-lfa-cp/Electrometer:103/fits/2025/06/
```

or top copy a particular file to local disk for further study:

```
mc cp lfa/rubinobs-lfa-cp/Electrometer:103/fits/2025/06/07/EM103_0_20250607_000389.fits .
```

All electrometer file from 2025-06-07 forward can be ingested with a command like:

```
butler ingest-photodiode main lsst.obs.lsst.LsstCam \
```



```
s3://lfa@rubinobs-lfa-cp/Electrometer:103/fits/2025/06/07/ \
--transfer=copy
```

This will automatically ingest into the LSSTCam/calib/photodiode collection, using the instrument class specified on the command line to set the collection prefix. The ingest process uses standard `lsst.resources.ResourcePath` objects to handle the path and files, which abstracts the S3 interface away, but do note the different syntax compared to the `mc` commands above. Also note that the trailing `/` in the path is required so that the files within the specified S3 directory are ingested.

A further complication is related to the SAL index for the electrometers; we have four electrometers (listed below in 5.1.1), three of which are connected to the main telescope. Re-iterating the constraint from above, only one photodiode measurement can be associated with an exposure, regardless of which electrometer did the measurement. Allowing multi-electrometer matching will need to be developed if we find a use case that requires that.

Name	Location	Model	Purpose
Electrometer:101	MainTel	Keysight	Part of CBP ( <b>CZW: needs clarification</b> )
Electrometer:102	MainTel	Keysight B2985B	Connected to output of CBP
Electrometer:103	MainTel	Keithley 6517B	Connected to LED and laser projector for flatfield screen
Electrometer:201	AuxTel	Keithley	Connected to lamps for the flatfield screen

### 5.1.2 Shutter Motion Profiles

Shutter motion profiles (SMPs) are similar data products, with a file written to the LFA that needs to be matched to the appropriate exposure. These data can be directly associated via the `obsId` key, which matches between exposure and SMP. Unlike the photodiode data, which can have only one per exposure, there are two SMPs per exposure, one for the shutter opening, and one for the shutter closing. The matching constraint still applies to these data, and so these two directions are stored in separate dataset types to prevent conflict.

## 6 Batch Process Submission (BPS) Templates

We have put together a set of convenient bps<sup>1</sup> template files that have been tested at USDF but should also be suitable testing at other locations. These templates are intended for batch processing of calibration construction and verification for both user collection testing and final calibration production. This section has a brief overview, but please see the README files at [https://github.com/lstt/cp\\_pipe/tree/main/bps/templates/README.md](https://github.com/lstt/cp_pipe/tree/main/bps/templates/README.md) and [https://github.com/lstt/cp\\_verify/tree/main/bps/templates/README.md](https://github.com/lstt/cp_verify/tree/main/bps/templates/README.md) for up-to-date information. Note that the templates assume that each calibration is certified into the overall calibration chain called `_${USER_CALIB_PREFIX}_${INSTRUMENT}calib_${TICKET}_${TAG}` between steps. A walkthrough example is in Section 7.

Use of the templates are controlled by a number of environment variables. At the writing of this documentation these include:

- `export USER_CALIB_PREFIX=""` or `export USER_CALIB_PREFIX=üerykoff`: Set to an empty string for official calibrations, or the user collection prefix with trailing slash.
- `export INSTRUMENT=LSSTComCam`: The name of the instrument.
- `export TICKET=DM-46562`: The name of the ticket associated with the calib construction.
- `export REPO=main`: The name of the butler repository to generate calibs.
- `export RAW_COLLECTION=LSSTComCamrawa11`: The name of the raw data collection.
- `export CALIB_COLLECTIONS=LSSTComCamcalibDM-46825`: Comma-separated list of curated or previously generated calibration collections to use as a starting point.
- `export TAG=newCalibs`: A human-readable tag to help indicate why a set of calibs were built (matched to the ticket number).
- `export RERUN=20250122a`: The rerun name to indicate when the calibrations were generated.
- `export BOOTSTRAP_RUN_NUMBER=1`: The bootstrap run number ensures that bootstrap run collection names are unique in case of an initial mistake.

---

<sup>1</sup><https://pipelines.lsst.io/modules/lstt.ctrl.bps.quickstart.html> and see also Gower et al. (2022) for background.

- `export SELECTION_BIAS=instrument='LSSTComCam'` and `selection_string`: The selection of raws to make the bootstrapBias and bias frames.
- `export SELECTION_DARK=instrument='LSSTComCam'` and `selection_string`: The selection of raws to make the bootstrapDark and dark frames.
- `export SELECTION_FLAT_BOOTSTRAP=instrument='LSSTComCam'` and `selection_string`: The selection of raws to make the bootstrapFlat.
- `export SELECTION_PTC=instrument='LSSTComCam'` and `selection_string`: The selection of raws to generate the PTC.
- `export SELECTION_PTC_LINEARIZER=$SELECTION_PTC`: The selection of raws to generate the linearizer; usually will be the same as the PTC selection.
- `export SELECTION_PTC_BFK=$SELECTION_PTC`: The selection of raws to generate the brighter-fatter kernel; usually will be the same as the PTC selection.
- `export SELECTION_PTC_CTI=$SELECTION_PTC`: The selection of raws to generate the charge-transfer-inefficiency dataset; usually will be the same as the PTC selection.
- `export SELECTION_FLAT_g=instrument='LSSTComCam'` and `selection_string`: The selection of raws to generate the g-band flat. There will need to be one selection for each ugrizy.

Additionally, for the verify templates we have:

- `export VERIFY_RERUN=20250122a`: The rerun name to indicate when the verification was run.
- `export SELECTION_BIAS_V=instrument='LSSTComCam'` and `selection_string`: The selection of raws to verify the bias frames.
- `export SELECTION_DARK_V=instrument='LSSTComCam'` and `selection_string`: The selection of raws to verify the dark frames.
- `export SELECTION_PTC_V=instrument='LSSTComCam'` and `selection_string`: The selection of raws to verify the PTC.
- `export SELECTION_PTC_LINEARIZER_V=$SELECTION_PTC_V`: The selection of raws to verify the linearizer; usually will be the same as the PTC selection.

- `export SELECTION_PTC_BFK_V=$SELECTION_PTC_V`: The selection of raws to verify the brighter-fatter kernel; usually will be the same as the PTC selection.
- `export SELECTION_FLAT_g_V=instrument='LSSTComCam and selection_string`: The selection of raws to verify the g-band flat. There will need to be one selection for each ugrizy.

## 7 Full Calibration Example

This section contains the commands used to generate, certify, chain, and verify calibrations for LSSTComCam DP1 processing. This is adapted from the script attached to DM-48520.

Note that this is not a complete construction of all calibrations; for this we used pre-existing linearity, ptc, and bfk.

In addition to the calibration building, this includes some checkpoints to download mosaic images for visual spot-checking to ensure that nothing has gone terribly wrong before continuing on all the way to verification reports.

```
export USER_CALIB_PREFIX=""
export INSTRUMENT=LSSTComCam
export TICKET=DM-48520
export REPO=main
export RAW_COLLECTION=LSSTComCam/raw/all
export CALIB_COLLECTIONS="LSSTComCam/calib/DM-48650,\
  LSSTComCam/calib/DM-47447/gainFixup/ptc.20241107a,\
  LSSTComCam/calib/DM-46360/isrTaskLSST/linearizer.20240926a,\
  LSSTComCam/calib/DM-46360/isrTaskLSST/bfk.20240926a,\
  LSSTComCam/calib/DM-46360/isrTaskLSST/ptc.20240926a"
export TAG=DP1
export RERUN=20250207a
export BOOTSTRAP_RUN_NUMBER=1

# These are certification environment variables.
export START_DATE=1970-01-01T00:00:00
export CERT_RERUN=20250207a

export SELECTION_BIAS="instrument='$INSTRUMENT' and \
  ((exposure.day_obs=20241208 and exposure.seq_num >= 696) or \
  (exposure.day_obs=20241206 and exposure.seq_num >=427)) and \
```

```

    exposure.observation_type='bias'"
export SELECTION_DARK="instrument='$INSTRUMENT' and \
    ((exposure.day_obs=20241208 and exposure.seq_num >= 696) or \
    (exposure.day_obs=20241206 and exposure.seq_num >=427)) and \
    exposure.observation_type='dark'"
export SELECTION_FLAT_BOOTSTRAP="instrument='$INSTRUMENT' and \
    exposure.day_obs=20240806 and exposure.science_program='BLOCK-T68' and \
    exposure.observation_reason='flat' and physical_filter='r_03'"

export SELECTION_PTC_CTI="instrument='$INSTRUMENT' and \
    exposure.day_obs=20240806 and \
    and exposure.science_program='BLOCK-T68' and \
    exposure.observation_reason='comcam_ptc' and \
    physical_filter='r_03'"

export SELECTION_FLAT_u="instrument='$INSTRUMENT' and \
    exposure.day_obs=20241110 and exposure.seq_num in (44..63)"
export SELECTION_FLAT_g="instrument='$INSTRUMENT' and \
    exposure.day_obs=20241126 and exposure.seq_num in (21..40)"
export SELECTION_FLAT_r="instrument='$INSTRUMENT' and \
    exposure.day_obs=20241111 and exposure.seq_num in (17..36)"
export SELECTION_FLAT_i="instrument='$INSTRUMENT' and \
    exposure.day_obs=20241120 and exposure.seq_num in (41..64) and exposure.seq_num != 58"
export SELECTION_FLAT_z="instrument='$INSTRUMENT' and \
    exposure.day_obs=20241112 and exposure.seq_num in (28..48)"
export SELECTION_FLAT_y="instrument='$INSTRUMENT' and \
    exposure.day_obs=20241119 and exposure.seq_num in (21..36)"

cd /sdf/data/rubin/user/erykoff/calibrations/LSSTComCam/DM-48520
mkdir -p images

# This is a quick check that the filled-in template looks
# correct (watch for missing env vars!)
cat $CP_PIPE_DIR/bps/templates/bps_biasBootstrap.yaml | envsubst

bps submit $CP_PIPE_DIR/bps/templates/bps_biasBootstrap.yaml

# Download a couple of images to check them.
butler retrieve-artifacts $REPO images \
    -d biasBootstrap \
    --collections ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/$TAG/biasBootstrapGen.$RERUN/run${BOOTSTRAP_RUN_NUMBER} \
    --where "instrument='${INSTRUMENT}' and detector in (5, 8)" --no-preserve-path

bps submit $CP_PIPE_DIR/bps/templates/bps_darkBootstrap.yaml

```

```

butler retrieve-artifacts $REPO images \
  -d darkBootstrap \
  --collections ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/$TAG/darkBootstrapGen.$RERUN/run${BOOTSTRAP_RUN_NUMBER} \
  --where "instrument='${INSTRUMENT}' and detector in (5, 8)" --no-preserve-path

bps submit $CP_PIPE_DIR/bps/templates/bps_flatBootstrap.yaml

bps submit $CP_PIPE_DIR/bps/templates/bps_defects.yaml

# We are certifying (applying a date range) and chaining the calibrations
# as we go, which makes downstream processing easier without having to
# track many collections.

butler certify-calibrations $REPO \
  --begin-date $START_DATE \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/${TAG}/defectGen.$RERUN \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/${TAG}/defects.$CERT_RERUN \
  defects

butler collection-chain $REPO \
  --mode redefine \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/${TAG}/defects.$CERT_RERUN

# The gainFixup PTC which is appropriate for the on-sky data after the
# ComCam warmup (when it was installed) is certified *after* the lab data.
# However, the lab PTC in LSSTComCam/calib/DM-46360/isrTaskLSST/ptc.20240926a
# is possibly more appropriate for the CTI calibration from the same
# dataset before the gains changed at the 0.3-0.5% level.

cat $CP_PIPE_DIR/bps/templates/bps_cti.yaml | envsubst

bps submit $CP_PIPE_DIR/bps/templates/bps_cti.yaml

butler certify-calibrations $REPO \
  --begin-date $START_DATE \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/${TAG}/ctiGen.$RERUN \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET/${TAG}/cti.$CERT_RERUN \
  cti

butler collection-chain $REPO \
  --mode prepend \
  ${USER_CALIB_PREFIX}$INSTRUMENT/calib/$TICKET \

```

```

    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/cti.${CERT_RERUN}

cat $CP_PIPE_DIR/bps/templates/bps_bias.yaml | envsubst

bps submit $CP_PIPE_DIR/bps/templates/bps_bias.yaml

butler retrieve-artifacts $REPO images \
    -d biasMosaic8 \
    --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/biasGen.${RERUN} \
    --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler certify-calibrations $REPO \
    --begin-date $START_DATE \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/biasGen.${RERUN} \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/bias.${CERT_RERUN} \
    bias

butler collection-chain $REPO \
    --mode prepend \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/bias.${CERT_RERUN}

cat $CP_PIPE_DIR/bps/templates/bps_dark.yaml | envsubst

bps submit $CP_PIPE_DIR/bps/templates/bps_dark.yaml

butler retrieve-artifacts $REPO images \
    -d darkMosaic8 \
    --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/darkGen.${RERUN} \
    --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler certify-calibrations $REPO \
    --begin-date $START_DATE \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/darkGen.${RERUN} \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/dark.${CERT_RERUN} \
    dark

butler collection-chain $REPO \
    --mode prepend \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \
    ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/dark.${CERT_RERUN}

cat $CP_PIPE_DIR/bps/templates/bps_flat_u.yaml | envsubst
```

```
bps submit $CP_PIPE_DIR/bps/templates/bps_flat_u.yaml

bps submit $CP_PIPE_DIR/bps/templates/bps_flat_g.yaml

bps submit $CP_PIPE_DIR/bps/templates/bps_flat_r.yaml

bps submit $CP_PIPE_DIR/bps/templates/bps_flat_i.yaml

bps submit $CP_PIPE_DIR/bps/templates/bps_flat_z.yaml

bps submit $CP_PIPE_DIR/bps/templates/bps_flat_y.yaml

butler retrieve-artifacts $REPO images \
  -d flat \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-u.${RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler retrieve-artifacts $REPO images \
  -d flat \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-g.${RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler retrieve-artifacts $REPO images \
  -d flat \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-r.${RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler retrieve-artifacts $REPO images \
  -d flat \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-i.${RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler retrieve-artifacts $REPO images \
  -d flat \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-z.${RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler retrieve-artifacts $REPO images \
  -d flat \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-y.${RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

butler certify-calibrations $REPO \
  --begin-date $START_DATE \
```



```
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-u.${RERUN} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-u.${CERT_RERUN} \  
 flat
```

```
butler collection-chain $REPO \  
 --mode prepend \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-u.${CERT_RERUN}
```

```
butler certify-calibrations $REPO \  
 --begin-date $START_DATE \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-g.${RERUN} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-g.${CERT_RERUN} \  
 flat
```

```
butler collection-chain $REPO \  
 --mode prepend \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-g.${CERT_RERUN}
```

```
butler certify-calibrations $REPO \  
 --begin-date $START_DATE \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-r.${RERUN} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-r.${CERT_RERUN} \  
 flat
```

```
butler collection-chain $REPO \  
 --mode prepend \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-r.${CERT_RERUN}
```

```
butler certify-calibrations $REPO \  
 --begin-date $START_DATE \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-i.${RERUN} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-i.${CERT_RERUN} \  
 flat
```

```
butler collection-chain $REPO \  
 --mode prepend \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \  
 ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-i.${CERT_RERUN}
```

```
butler certify-calibrations $REPO \  
 --begin-date $START_DATE \  
 
```

```

${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-z.${RERUN} \
${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-z.${CERT_RERUN} \
flat

butler collection-chain $REPO \
  --mode prepend \
  ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \
  ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-z.${CERT_RERUN}

butler certify-calibrations $REPO \
  --begin-date $START_DATE \
  ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-y.${RERUN} \
  ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-y.${CERT_RERUN} \
  flat

butler collection-chain $REPO \
  --mode prepend \
  ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET} \
  ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flat-y.${CERT_RERUN}

#####
## Verification section.
#####

export VERIFY_RERUN=20240210a
export SELECTION_SCIENCE_V="instrument='$INSTRUMENT' and \
  exposure in (2024120100160,2024120700307,2024120700297,2024120300156,\
  2024120700339,2024112000223)"
# Other exports
export SELECTION_BIAS_V="instrument='$INSTRUMENT' and \
  exposure in (2024102100509,2024102600031,2024102900175,2024110100164,\
  2024110800008,2024111200387,2024111600316,2024112200390,2024112300315,\
  2024112400006,2024112700443,2024120500004,2024120500266,2024120500272,\
  2024120500291,2024120700571,2024120700576,2024120700592,2024120700598,\
  2024120800704,2024120800724,2024120900006,2024120900284,2024121000596,\
  2024121100007)"
export SELECTION_DARK_V="instrument='$INSTRUMENT' and \
  exposure in (2024102300019,2024102300035,2024102500019,2024102500044,\
  2024103000012,2024103000037,2024110100149,2024110100175,2024110500279,\
  2024110500301,2024111100307,2024111100334,2024111600322,2024111600350,\
  2024112000330,2024112000355,2024112700437,2024112700455,2024120100481,\
  2024120100507,2024120700582,2024120700588,2024120700604,2024120700610,\
  2024121000007,2024121000590,2024121000612)"
export SELECTION_FLAT_u_V=$SELECTION_FLAT_u

```

```
export SELECTION_FLAT_g_V=$SELECTION_FLAT_g
export SELECTION_FLAT_r_V=$SELECTION_FLAT_r
export SELECTION_FLAT_i_V=$SELECTION_FLAT_i
export SELECTION_FLAT_z_V=$SELECTION_FLAT_z
export SELECTION_FLAT_y_V=$SELECTION_FLAT_y

# Run science frames.

cat $CP_VERIFY_DIR/bps/templates/bps_verify_science.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_science.yaml

butler retrieve-artifacts $REPO images \
  -d verifyScience8 \
  --collections ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyScience.${VERIFY_RERUN} \
  --where "instrument='${INSTRUMENT}'" --no-preserve-path --find-first

# Other verifications.

cat $CP_VERIFY_DIR/bps/templates/bps_verify_bias.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_bias.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_dark.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_dark.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_flat_u.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_flat_u.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_flat_g.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_flat_g.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_flat_r.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_flat_r.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_flat_i.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_flat_i.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_flat_z.yaml | envsubst
```

```
bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_flat_z.yaml

cat $CP_VERIFY_DIR/bps/templates/bps_verify_flat_y.yaml | envsubst

bps submit $CP_VERIFY_DIR/bps/templates/bps_verify_flat_y.yaml

# Make TAXICAB reports.
# Note that the order matters, and should be verify runs first,
# followed by gen runs.

$CP_VERIFY_DIR/bin/cpv_report.py -r $REPO \
  -O ~/public_html/cpv_reports/TAXICAB-23 \
  --do_overwrite \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyBias.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyDark.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyFlat-u.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyFlat-g.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyFlat-r.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyFlat-i.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyFlat-z.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/verifyFlat-y.${VERIFY_RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/biasGen.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/darkGen.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-u.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-g.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-r.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-i.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-z.${RERUN} \
  -c ${USER_CALIB_PREFIX}${INSTRUMENT}/calib/${TICKET}/${TAG}/flatGen-y.${RERUN}
```

## 8 Conclusions

### A References

Gower, M., Kowalik, M., Lust, N.B., Bosch, J.F., Jenness, T., 2022, arXiv e-prints, arXiv:2211.15795

(arXiv:2211.15795), doi:10.48550/arXiv.2211.15795, ADS Link

**[RTN-075]**, Guy, L., 2024, *Charge to the Rubin Telescope and Auxiliary Instrumentation Calibration Acceptance Board*, Technical Note RTN-075, NSF-DOE Vera C. Rubin Observatory, URL <https://rtn-075.lsst.io/>

## B Acronyms

Acronym	Description
BPS	Batch Production Service
BTS	Base (La Serena) Test Stand
CBP	Collimated Beam Projector
CP	Contingency Planning
CTI	Charge Transfer Inefficiency
DM	Data Management
DMTN	DM Technical Note
DP1	Data Preview 1
FGCM	Forward Global Calibration Method
HTML	HyperText Markup Language
ISO	Information Security Officer
LATISS	LSST Atmospheric Transmission Imager and Slitless Spectrograph
LED	Light-Emitting Diode
LFA	Large File Annex
LOVE	LSST Operators Visualization Environment
LSSTCam	LSST Science Camera
LSSTComCam	Rubin Commissioning Camera
PTC	Photon Transfer Curve
RTN	Rubin Technical Note
S3	(Amazon) Simple Storage Service
SAL	Service Abstraction Layer
SLAC	SLAC National Accelerator Laboratory
TAI	International Atomic Time
TAXICAB	Telescope and Auxiliary Instrumentation Calibration Acceptance Board
TBD	To Be Defined (Determined)

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TTS	Tucson Test Stand
US	United States
USDF	United States Data Facility
bps	bit(s) per second

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